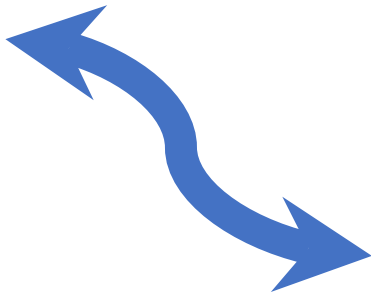


# A Multi-Wavelength Photometer for Small Satellite Heliophysics

**We Sort Light**



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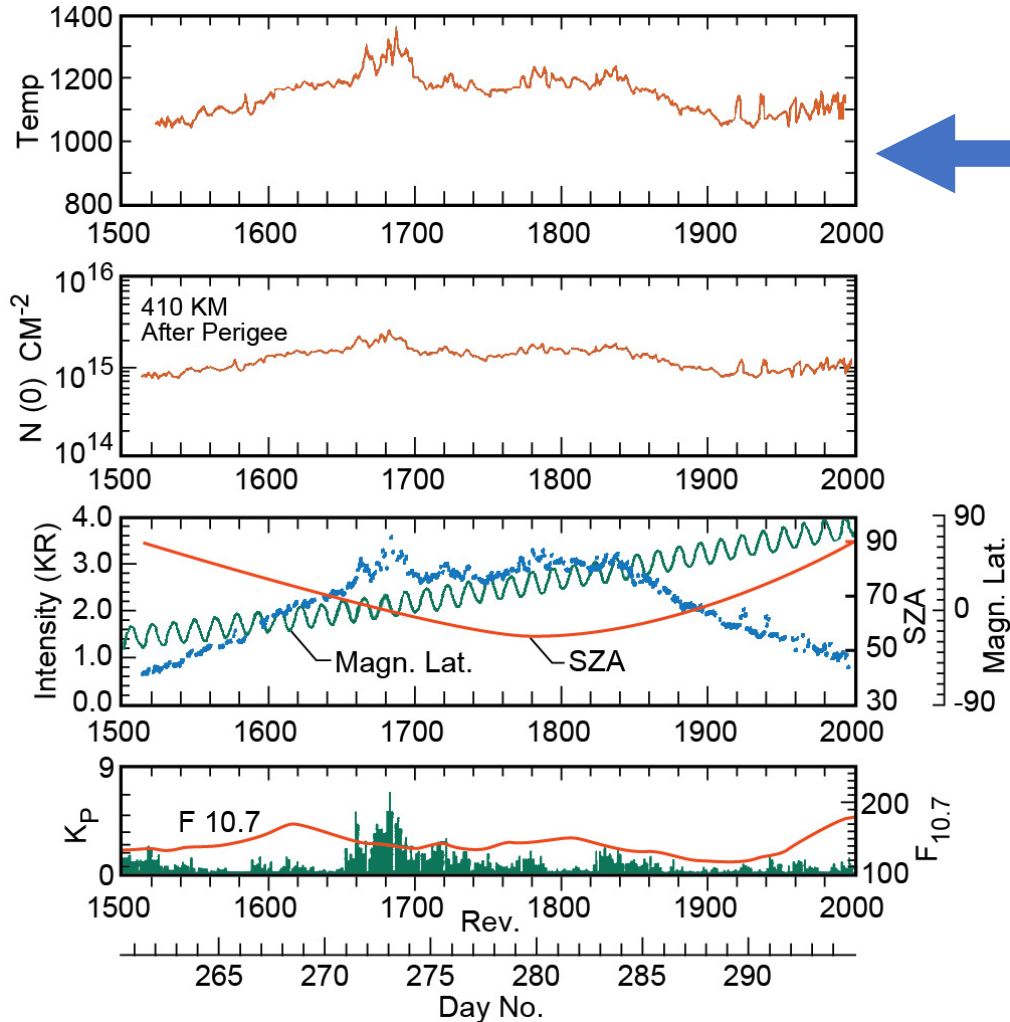


COMPUTATIONAL  
PHYSICS, INC.



# Motivation

**Small-Sat VUV sensor to compliment GUVI, SUSI, GOLD & other larger missions.**



OGO-6 (1976) zenith OI 130.4 nm photometer data and derived products N(O) and Texo

**Strickland and Thomas 1976**

# Motivation

## Characterize thermospheric variability

Proportionality of  $O/N_2 \propto 135.6/LBH$

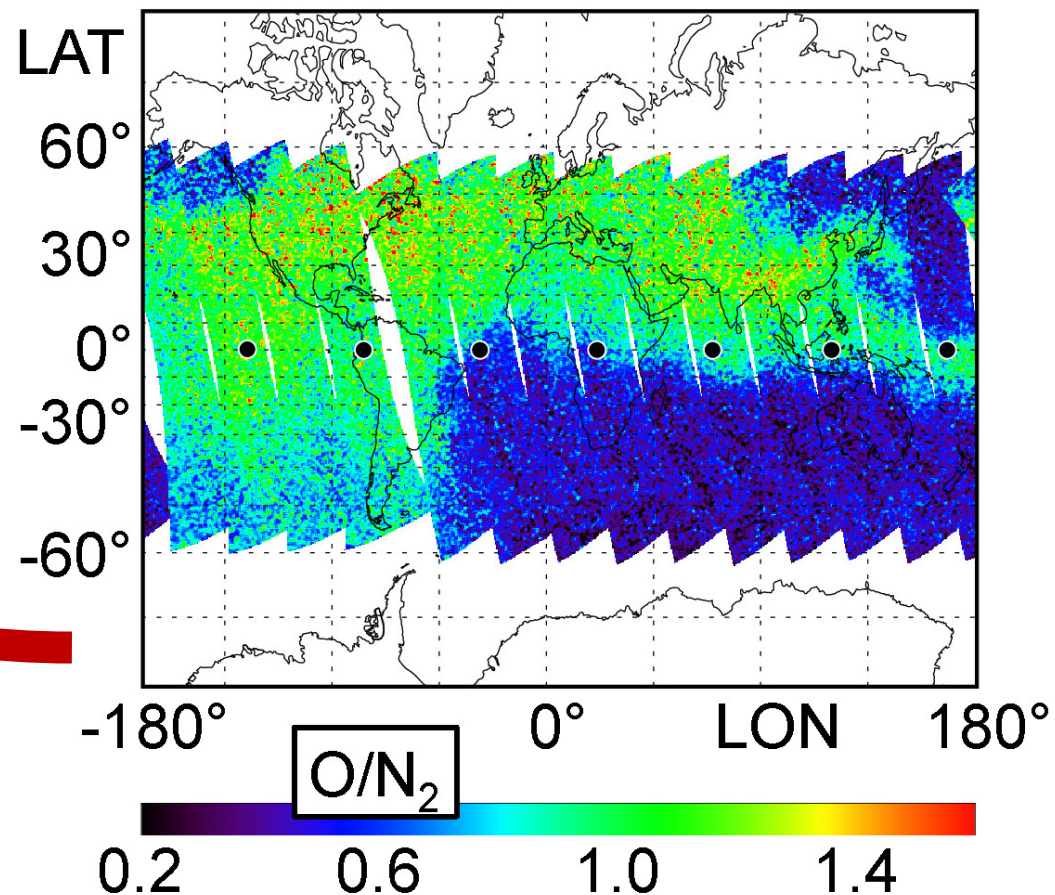
**O/N<sub>2</sub> variability between quiet and disturbed conditions.**

Strickland et al. [1995]  
Evans et al. [1995]  
Strickland et al. [1999b; 2001]  
Daniell and Strickland [2001]  
Meier et al. [2005]  
Crowley et al. [2006]

**Storm time O/N<sub>2</sub>**



GUVI O/N<sub>2</sub> 21 Nov 2003



# Designed for Targeted Observations

	SCIENCE TARGET			
Wavelength	Dayside O/N <sub>2</sub>	Auroral Energetics		Nightside F-Region
OI 130.4 nm				X
OI 135.6 nm	X	X		X
LBHS 150 nm			X	
LBHL 170 nm	X	X	X	

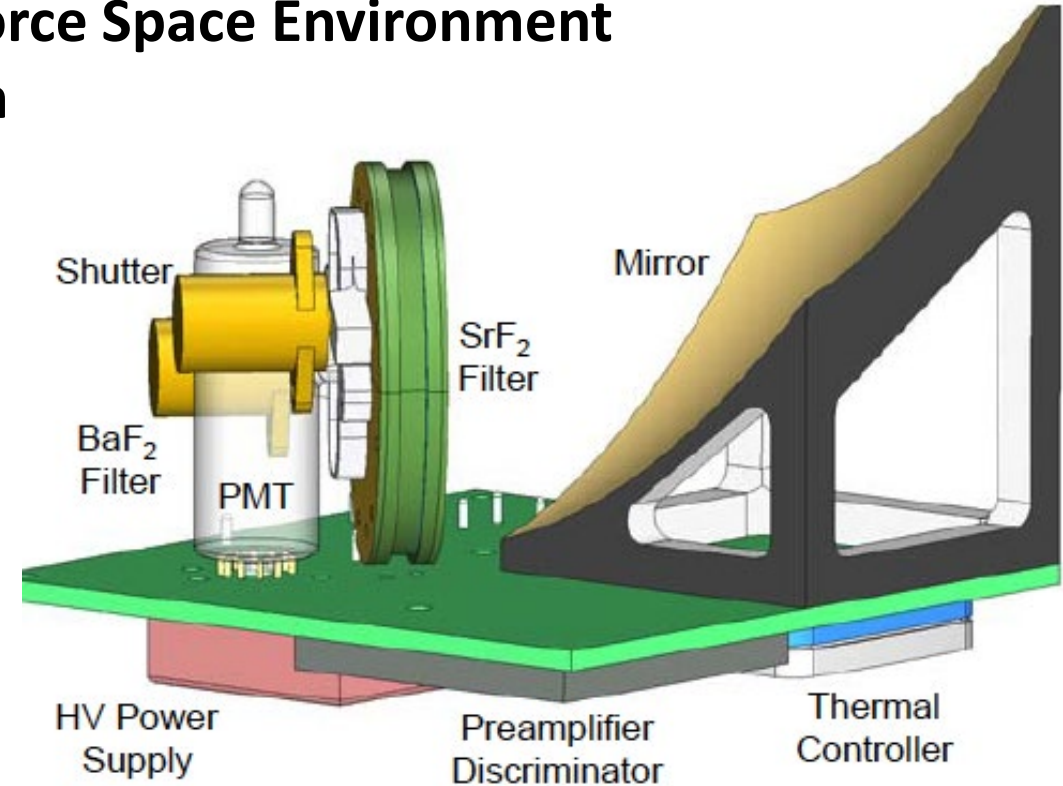
**ALSO**



**Lyman Alpha at 121.6**

# CubeSat Tiny Ionospheric Photometer (CTIP)

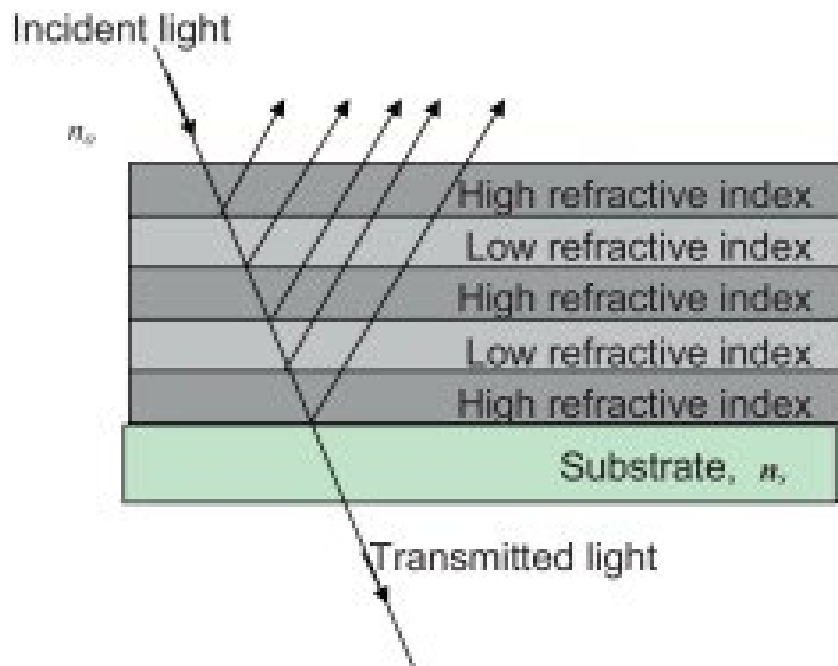
- **O<sup>+</sup> density by 135.6-nm photoemission (nighttime)**
- **Similar to Tiny Ionospheric Photometer on COSMIC**
- **Launched November 2013 on the US Air Force Space Environment Nano-Satellite Experiment (SENSE) mission**





# Reflective filters for the VUV

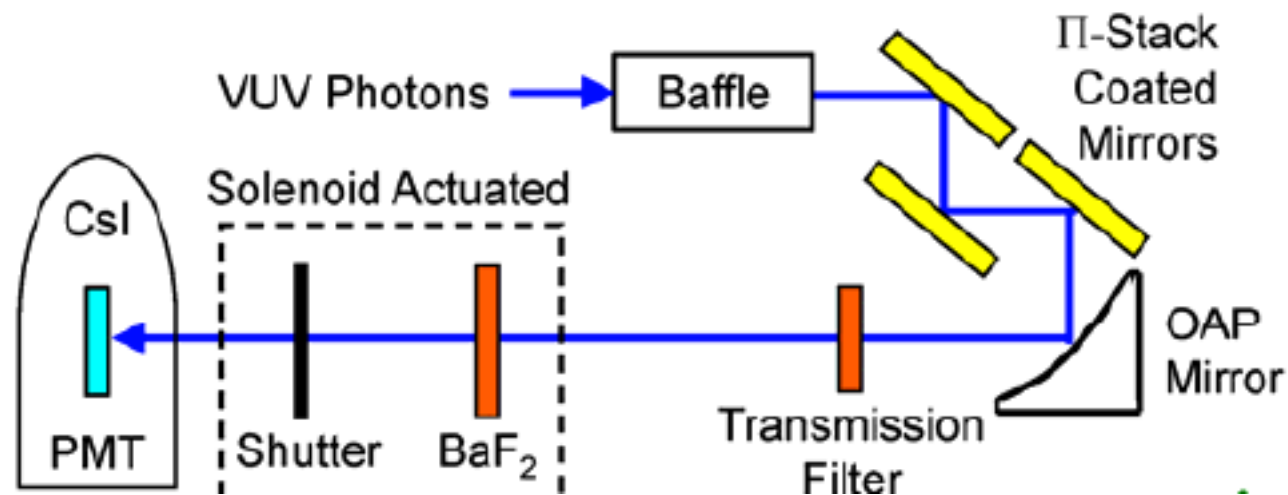
Pioneered by Torr and Zukic for Polar UVI



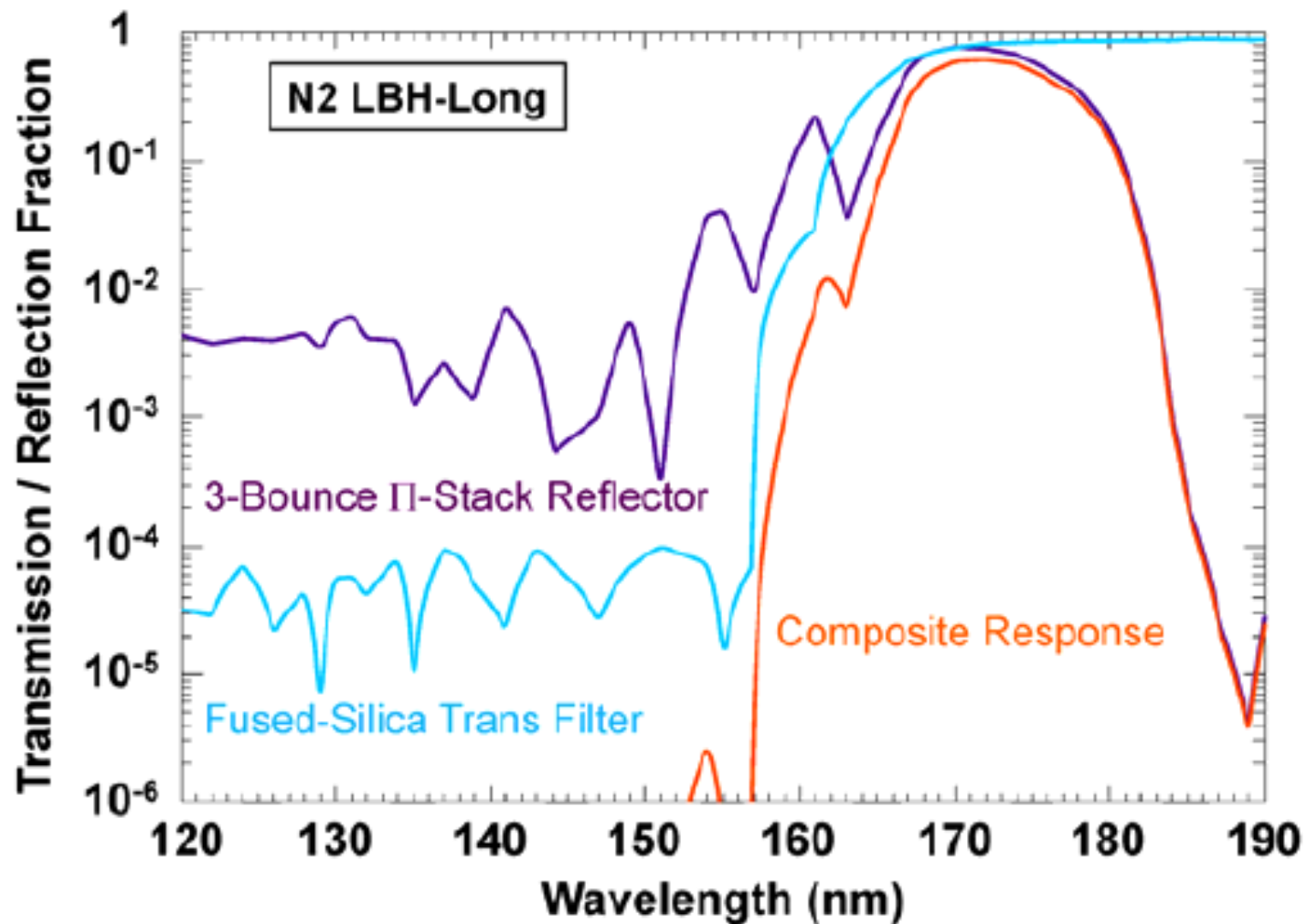
$H+L=0.5\lambda$   
 $\Pi$  phase shift

One mirror not sufficient;  
need multiple bounces for  
background rejection.

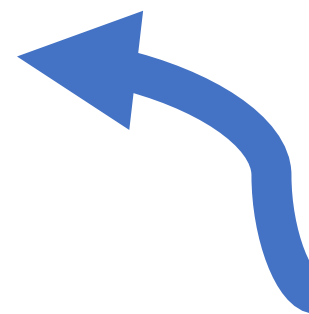
$\Pi$  multilayer stacks,  
alternating layers of high  
and low index material  
~30 layers



# Reflective filters for the VUV

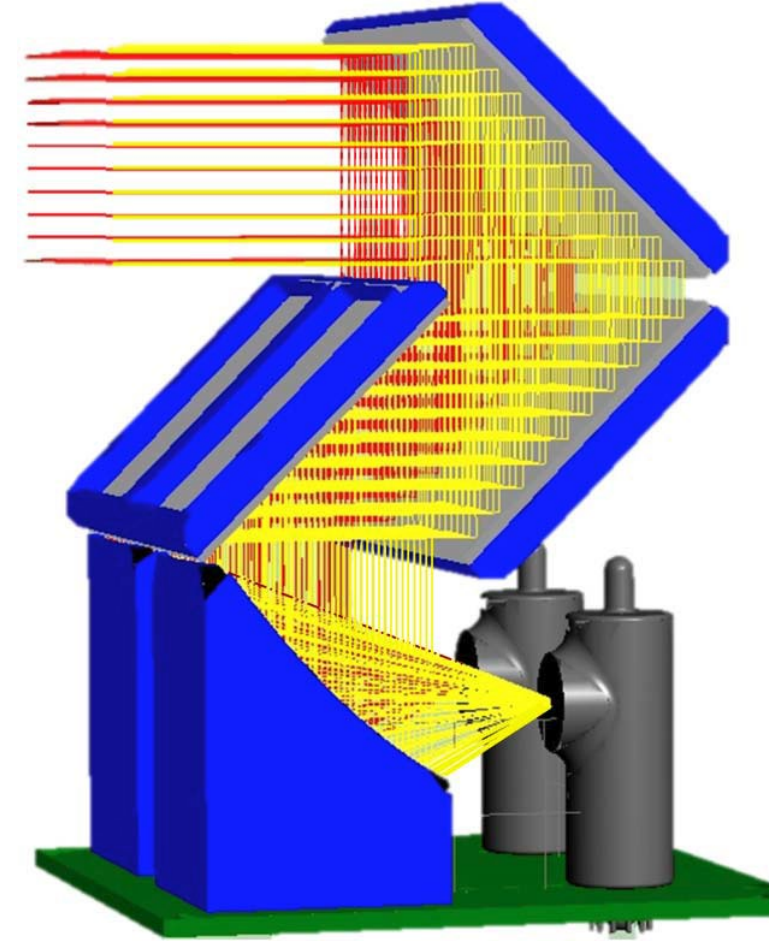
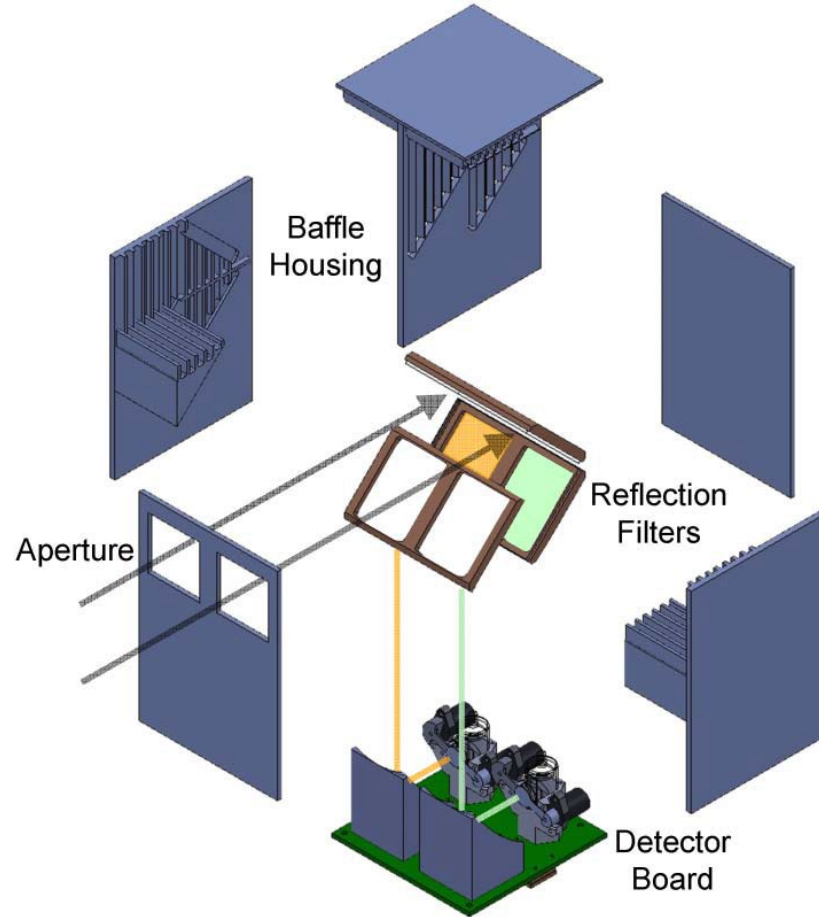
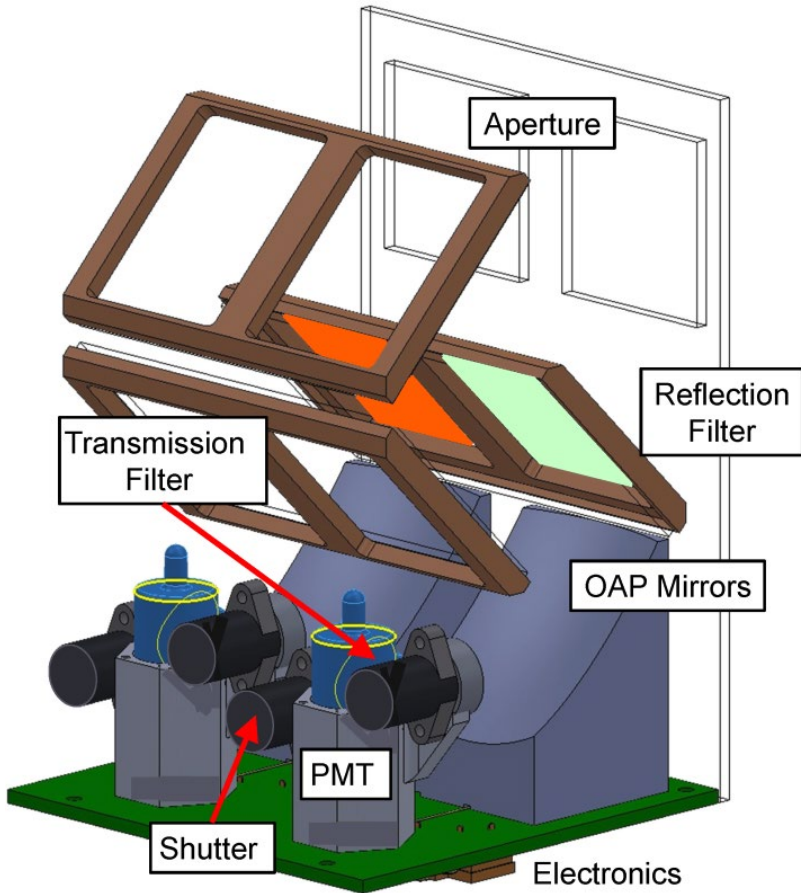


Zukic et al., 1990a  
Zukic et al., 1990b  
Torr et al., 1993



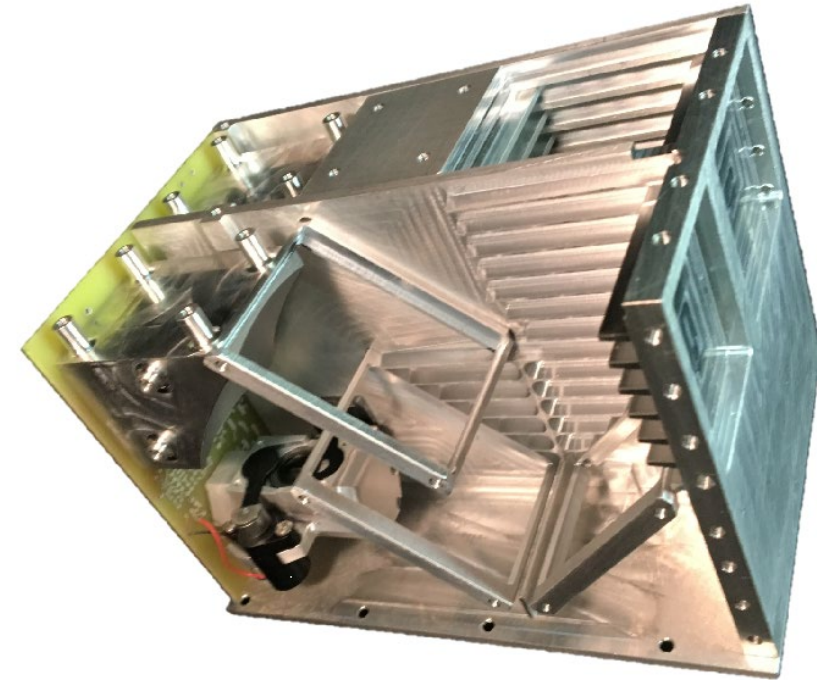
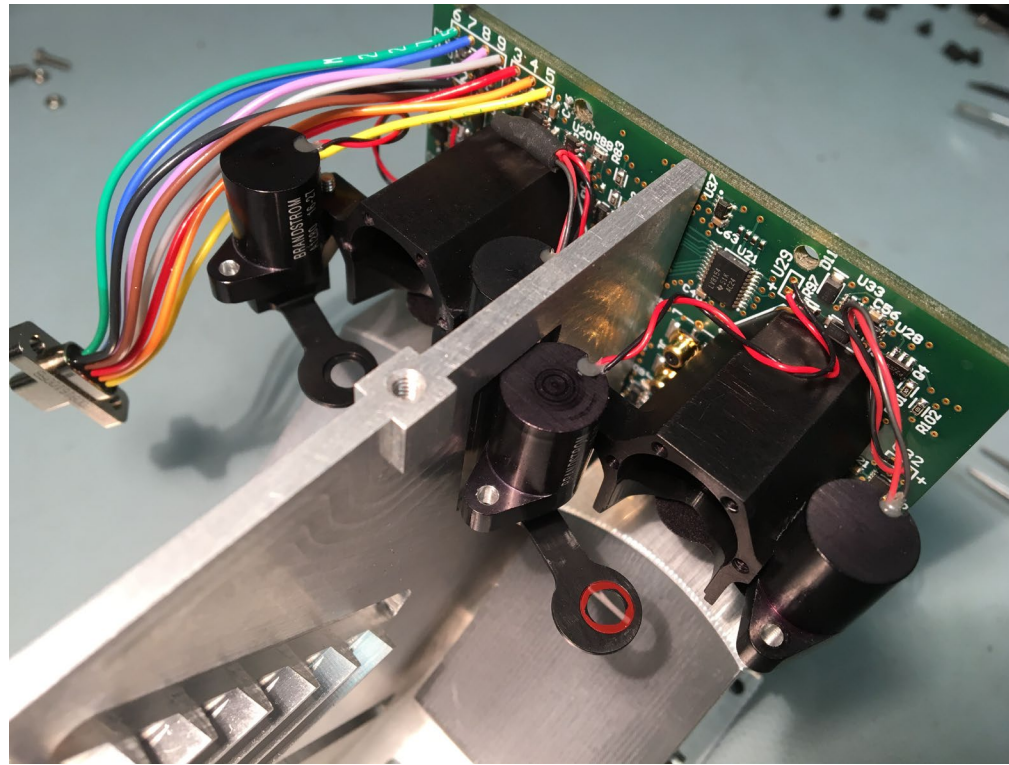
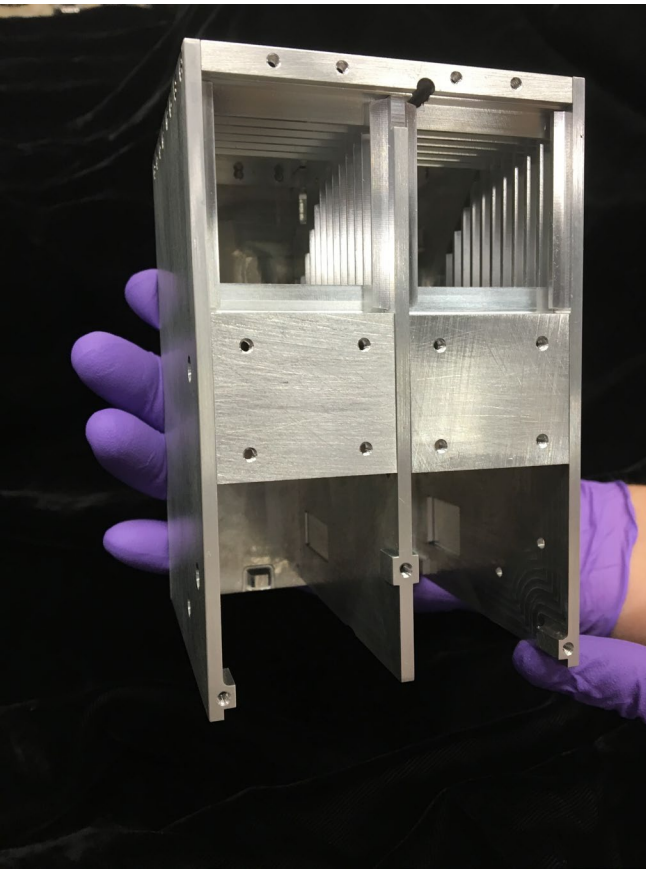
Expected LBH-Long  
composite filter response

# Two Wavelength Photometer The Design



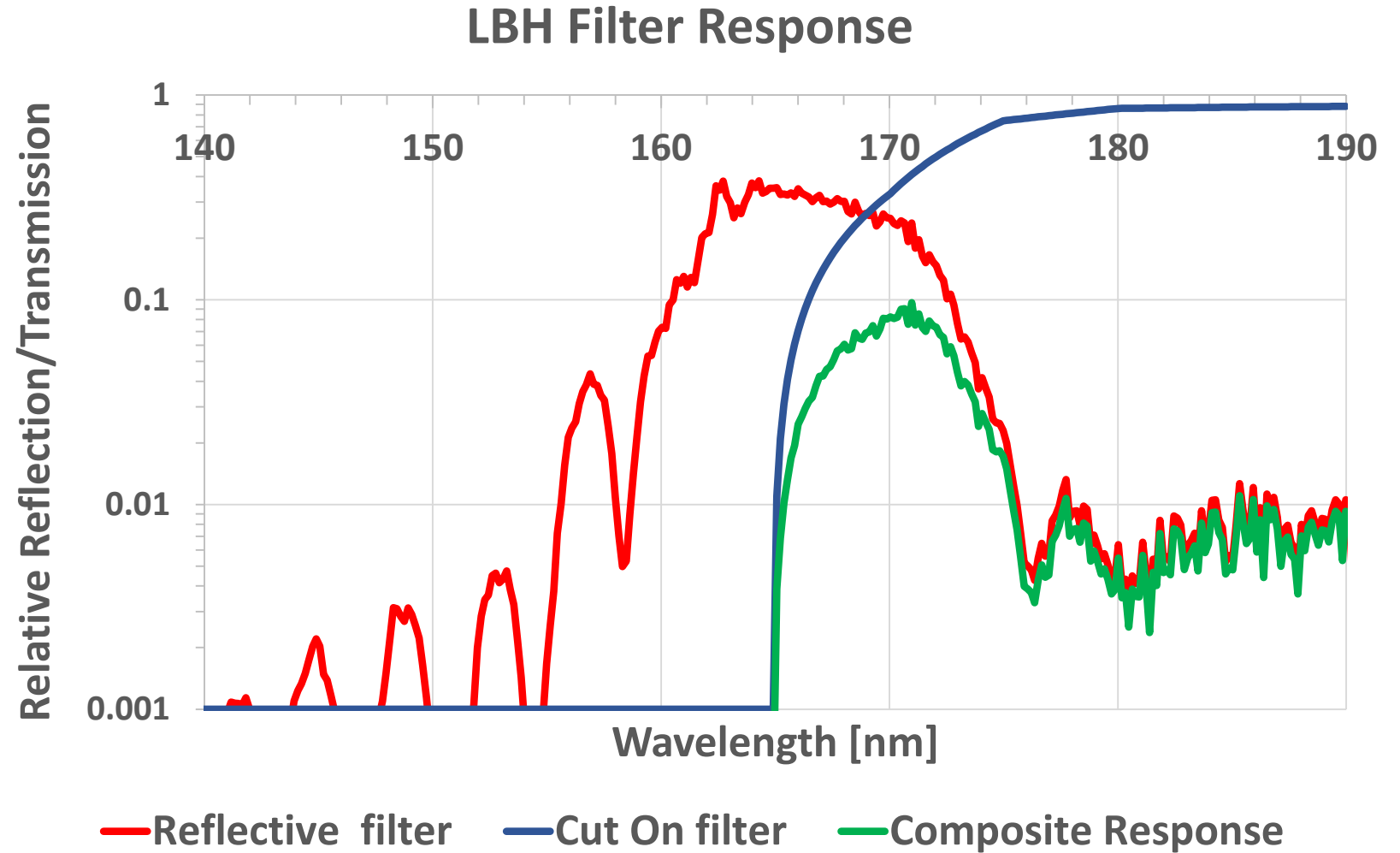


# Two Wavelength Photometer: The Build



# Results

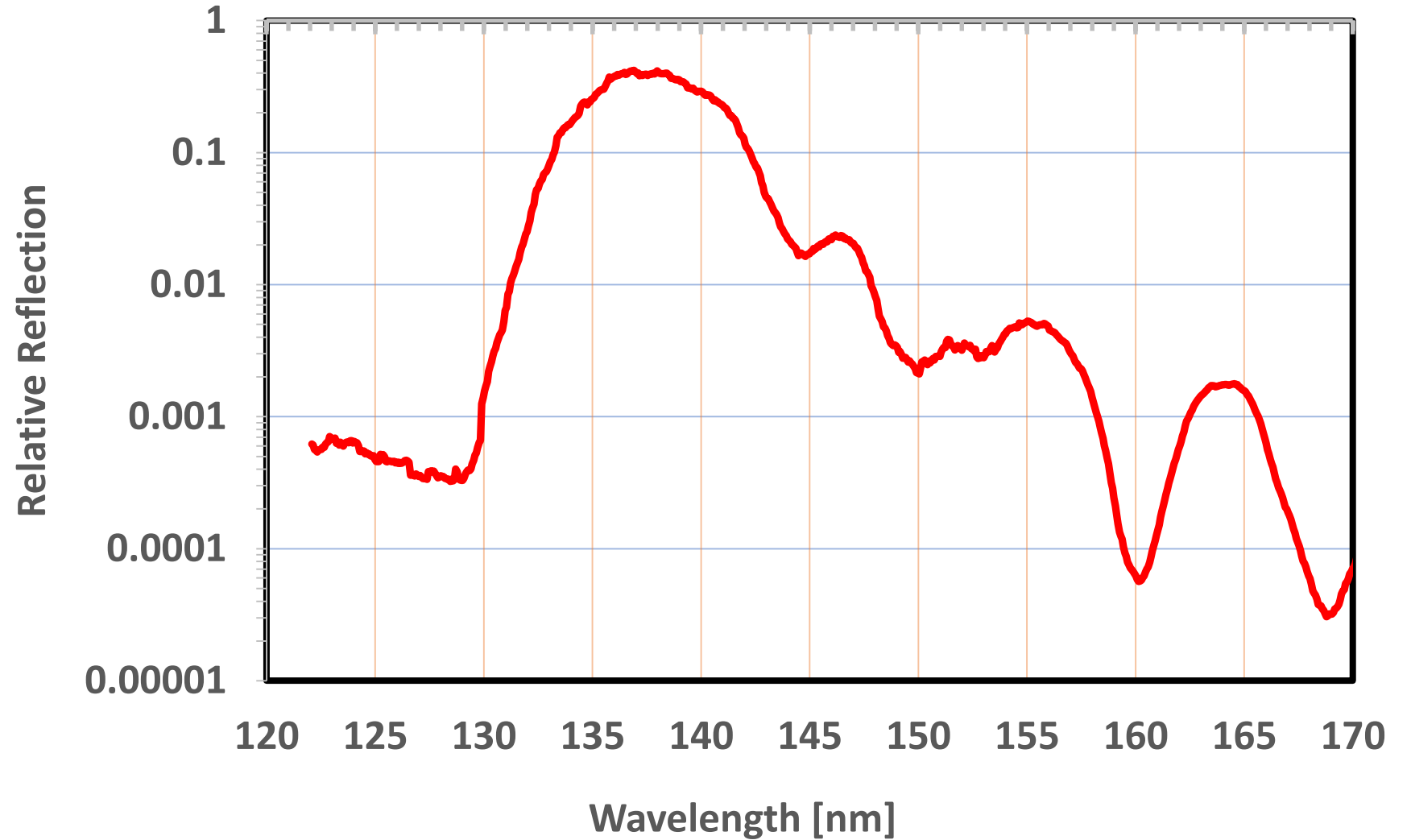
- ✓ FWHM ~ 9nm
- ✓ Max T of 10%



# Results

- ✓ Max T of 55%
- ✓ FWHM ~ 8nm

## Spectrum of 135.6nm Reflection Filter



Parameter	OI Channel	N <sub>2</sub> Channel
Center Wavelength	135.6 nm	170.0 nm
Bandwidth	5.0 nm	9.0 nm
Baseline Responsivity <sup>1</sup>	146 counts/R-s	340 counts/R-s
Detector Noise	< 10 counts/R-s	
Min Signal (Night / Day) <sup>2</sup>	2.7 R / 1590 R	NA / 680 R
Max Signal (Night / Day) <sup>2</sup>	213 kR / 5100 kR	NA / 2550 kR
Field of View	6.5° X 6.5°	
Nightside Resolution <sup>3</sup>	34 km X 34 km	NA
Auroral Resolution <sup>3</sup>	56 km X 56 km	56 km X 56 km
Orbit Averaged Power <sup>4</sup>	1 W	
Survival Temperature	-24° to +61°C	
Operating Temperature	-13° to +55°C	
Mass (Margin)	800 g (80 g)	
Volume	1400 cm <sup>3</sup>	
Form Factor	9.5 cm X 9.5 cm X 12 cm	
Power & Communication	5 ± 0.2 V <sub>DC</sub> , RS422 Serial	
CTIP Software Reuse	85%	

# The End

The authors would like to thank NASA Heliophysics for funding this project through the HTiDES program under grant NNX15AK45G